Determination of Neutrino Mass Hierarchy from T2K and NOvA: Neutrino Running ONLY

Stephen Parke - Fermilab with Olga Mena and Hiroshi Nunokawa hep-ph/0609011

many discussions with Hisakazu Minakata HM, NN & SP: Phys. Rev. D68, 013010 (2003)

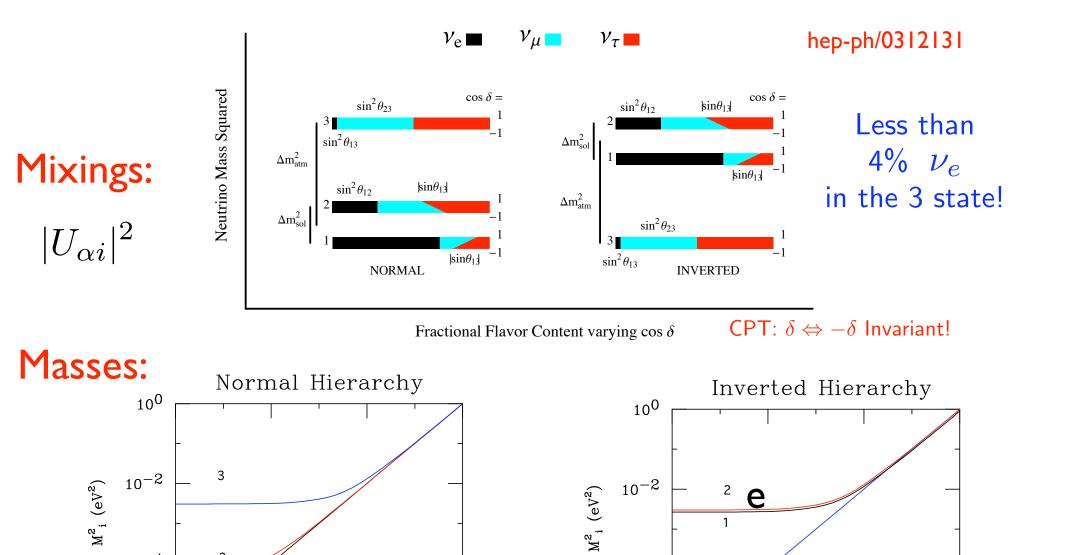
Outline:

- Introduction $P(\nu_{\mu} \rightarrow \nu_{e})$
- T2K and NOvA
- LBL: Ways to determine Neutrino Mass Hierarchy
- Further out ways
- Conclusions

Credits:

For T2K: Aihara P5 talk

For NOvA: Karol Lang NO-VE 2007 talk



 10^{-4}

 10^{-6}

 10^{-6}

10-2

 ${\rm M^2_{\ lite}\ (eV^2)}$

100

 10^{-4}



 10^{-4}

 10^{-6}

 10^{-6}

3

 10^{-4}

10-2

 $M^2_{lite} (eV^2)$

100

Long Baseline $u_{\mu} \rightarrow \nu_{e} \text{ or } \nu_{e} \rightarrow \nu_{\mu}$

- SUPERBEAMS: (0.4 to 4 MW)
 - Counting Expts (3 ways)
 - Spectrum Measurement
- NEW NEUTRINO BEAMS
 - Neutrino Factory (muon storage ring)
 - High Gamma Beta Beams

Vacuum LBL:

$$\nu_{\mu} \rightarrow \nu_{e}$$

amplitude:

$$\mathcal{A}(\nu_{\mu} \to \nu_{e}) = U_{\mu 3}^{*} U_{e3} e^{-im_{3}^{2}L/2E} + U_{\mu 2}^{*} U_{e2} e^{-im_{2}^{2}L/2E} + U_{\mu 1}^{*} U_{e1} e^{-im_{1}^{2}L/2E}$$

eliminate last term using unitarity:

$$U_{\mu 1}^* U_{e1} + U_{\mu 2}^* U_{e2} + U_{\mu 3}^* U_{e3} = 0$$

$$\mathcal{A}(\nu_{\mu} \to \nu_{e}) = U_{\mu 3}^{*} U_{e 3} \sin \Delta_{31} e^{-i\Delta_{32}} + U_{\mu 2}^{*} U_{e 2} \sin \Delta_{21}$$

and $\Delta_{ij} \equiv \delta m_{ij}^2 L/4E$ is the kinematical phase:

$$\nu_{\mu} \rightarrow \nu_{e}$$

U=MNS matrix

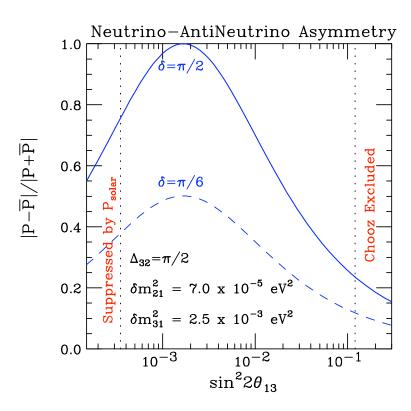
$$P_{\mu o e}pprox |\sqrt{P_{atm}}e^{-i(\Delta_{32}\pm\delta)}+\sqrt{P_{sol}}|^2$$
 $\Delta_{ij}=|\delta m_{ij}^2|L/4E$ CP violation !!!

where
$$\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \sin \Delta_{31}$$

and
$$\sqrt{P_{sol}} = \cos \theta_{23} \sin 2\theta_{12} \sin \Delta_{21}$$

$$P_{\mu
ightarrow e} pprox \mid \sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}} \mid^2$$

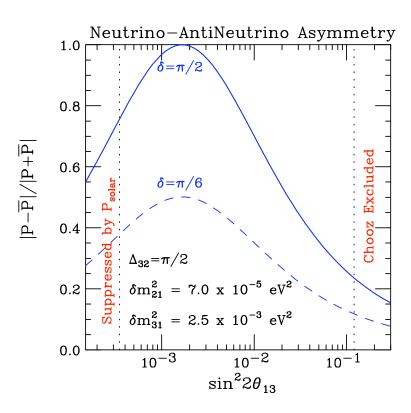
Asymmetry Peaks:



$$\sqrt{P_{atm}} = \sqrt{P_{sol}}$$

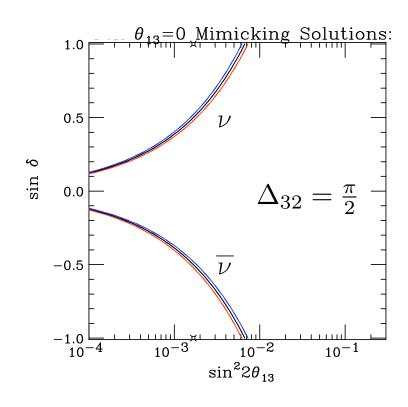
$$P_{\mu \to e} pprox |\sqrt{P_{atm}} e^{-i(\Delta_{32} \pm \delta)} + \sqrt{P_{sol}}|^2$$

Asymmetry Peaks:



$$\sqrt{P_{atm}} = \sqrt{P_{sol}}$$

Zero Mimicking Solutions: $P_{\mu \rightarrow e} = P_{sol}$



$$\sqrt{P_{atm}} = -2\sqrt{P_{sol}}\cos(\Delta_{32} \pm \delta)$$

$$P_{\mu o e} pprox |\sqrt{P_{atm}}e^{-i(\Delta_{32}\pm\delta)}+\sqrt{P_{sol}}|^2$$
 with $N^2\sin 2 heta$ is invariant

 $\{\delta m^2\sin2\theta\}$ is invariant

where
$$\sqrt{P_{atm}} = \sin \theta_{23} \sin 2\theta_{13} \left(\frac{\sin(\Delta_{31} \mp aL)}{(\Delta_{31} \mp aL)} \Delta_{31} \right)$$

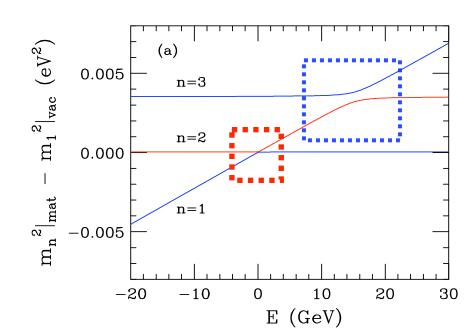
and
$$\sqrt{P_{sol}} = \cos\theta_{23}\sin2\theta_{12}\frac{\sin(aL)}{(aL)}\Delta_{21}$$

in vac $\sin \Delta_{21}$

in vac $\sin \Delta_{31}$

$$a = G_F N_e / \sqrt{2} = (4000 \ km)^{-1}$$
,

$$\pm = sign(\delta m^2_{31}) \quad \Delta_{ij} = |\delta m^2_{ij}|L/4E$$



Counting Expts near First Osc. Max.

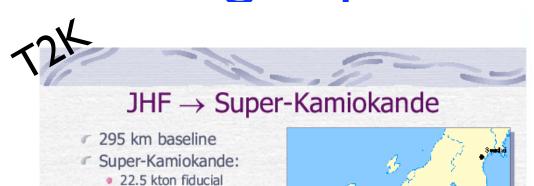
 $\mathsf{JHF} \to \mathsf{Super\text{-}Kamiokande}$

- 295 km baseline
- Super-Kamiokande:
 - 22.5 kton fiducial
 - Excellent e/μ ID
 - Additional π⁰/e ID
- Hyper-Kamiokande
 - 20× fiducial mass of SuperK
- Matter effects small
- Study using fully simulated and reconstructed data

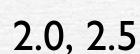


2.0, 2.5 or 3.0 deg

Counting Expts near First Osc. Max.

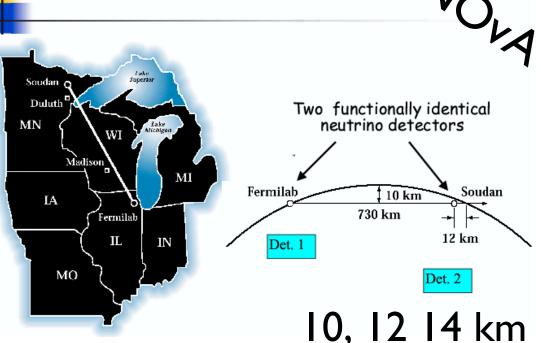


- Excellent e/μ ID
- Additional π⁰/e ID
- Hyper-Kamiokande
 - 20× fiducial mass of SuperK
- Matter effects small
- Study using fully simulated and reconstructed data



Kamiokande 295km

The NUMI Beamline



Water Cherenkov detector

● 1996.4 Start data taking (SK-I)

39m

- 2001.7 Stop data taking for detector upgrade
- 2001.11 Accident (6777 inner PMTs, 1100 outer PMTs were destroyed)

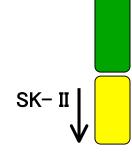
partial reconstruction of the detector

- 2002.10 Resume data taking (SK- II) mostly for K2K (photocathode coverage of 20%, 7MeV
- 2005.10 Start full recovery work

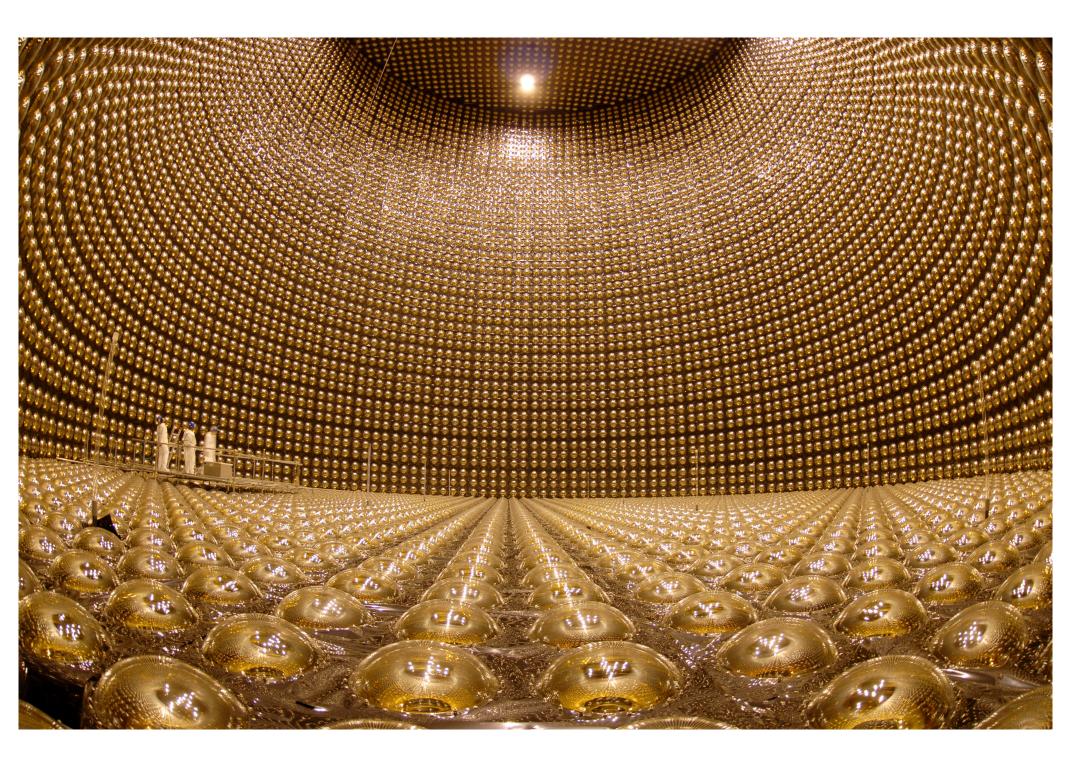
- 2700 w.e. overbuden
- 50,000 ton (22,500 ton fid.)
- 11,146 20 inch PMTs
- Photo cathode coverage: 40% of surface
- 1,885 anti-counter PMTs



Acrylic + FRP vessel



SK-I
[~]5years





NOvA Far Detector we would like to build



□TAD = Totally Active Detector
PVC = passive material

□ mass N kT (N large) ~80% scintillator

~20% PVC extrusions

■Modular structure32 cells/extrusion12 extrusions/plane

1984 planes

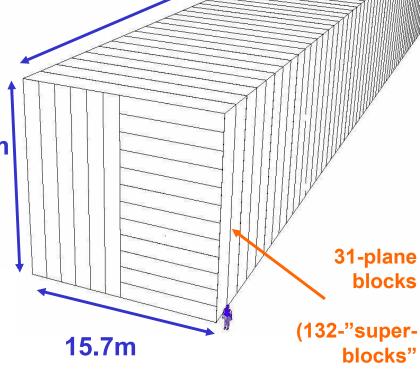
□Cell dimensions:

3.9 cm x 6 cm x 15.7m

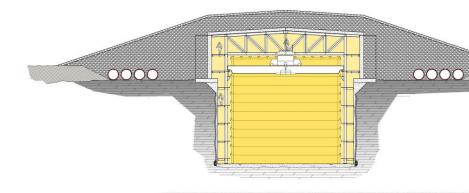
□U-shaped 0.7 mm WLS fiber into APD



15.7m

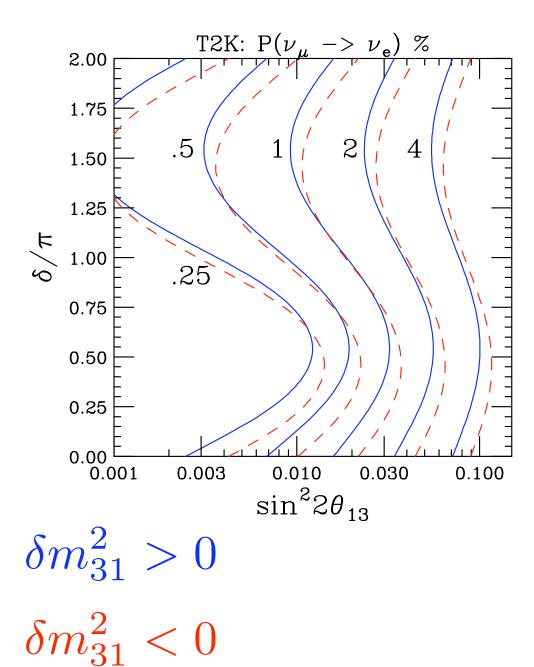


We will build as much of this as the funding will allow..



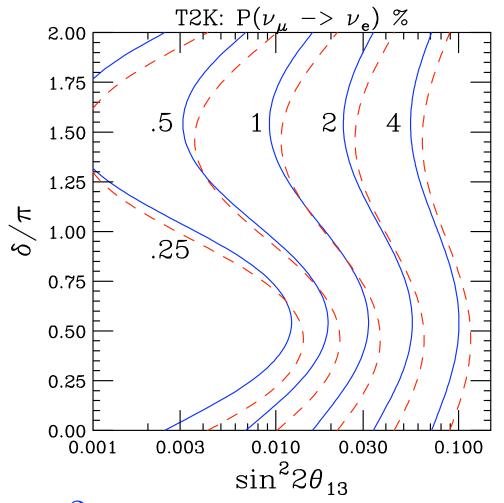
Far Detector

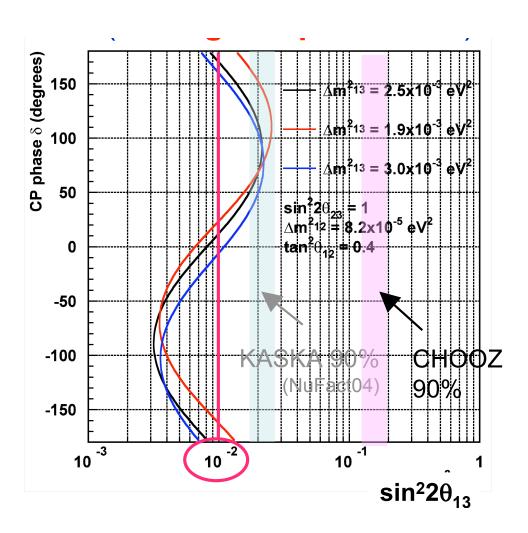
T2K:



T2K:

Aihara for T2K, P5 talk



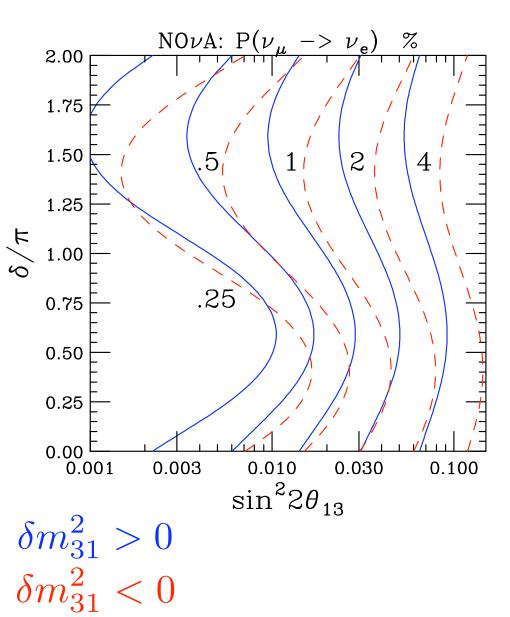


$$\delta m_{31}^2 > 0$$

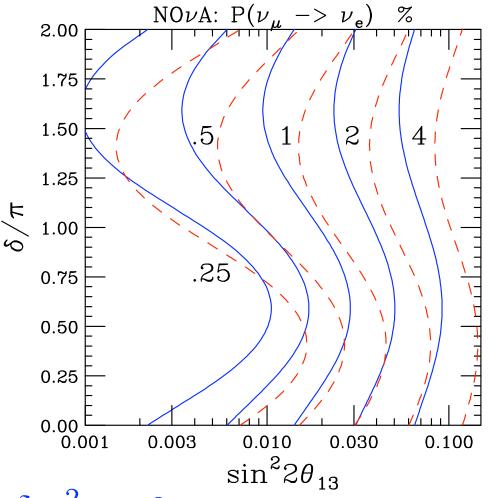
$$\delta m_{31}^2 > 0$$
 $\delta m_{31}^2 < 0$

Phase I Sensitivity approx 0.5%

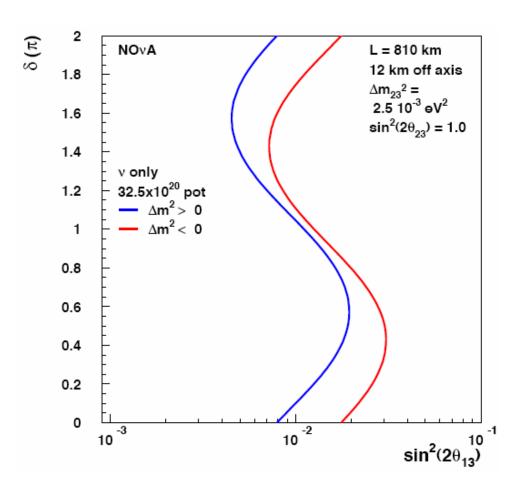
NOvA:



NOvA:



NOvA @ NO-VE 2007

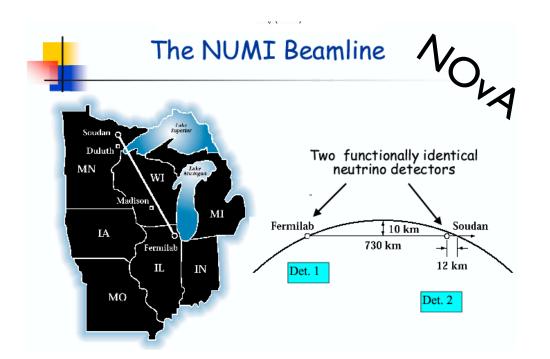


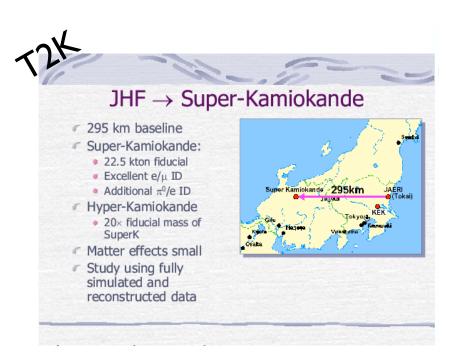
 $\delta m_{31}^2 > 0$ $\delta m_{31}^2 < 0$ Phase I Sensitivity approx 0.5%

5 years with ν only run

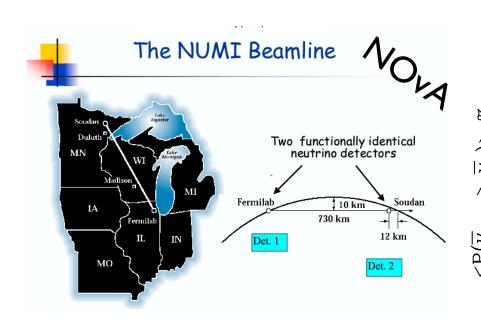
Counting Expts at First Osc. Max.

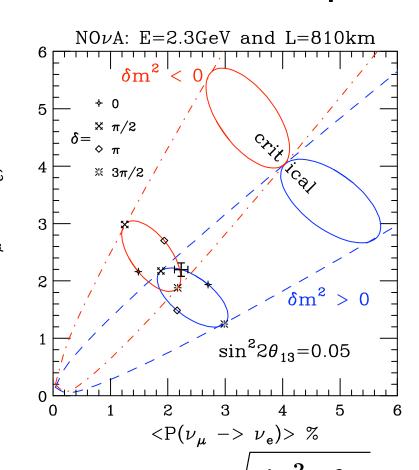
- Neutrino v Anti-Neutrino One Expt.
- Neutrino v Neutrino Two Expts Different L's and EQUAL E/L's
- Neutrino v Anti-Neutrino Two Expts Different L's





Neutrino v Anti-Neutrino One Expt.





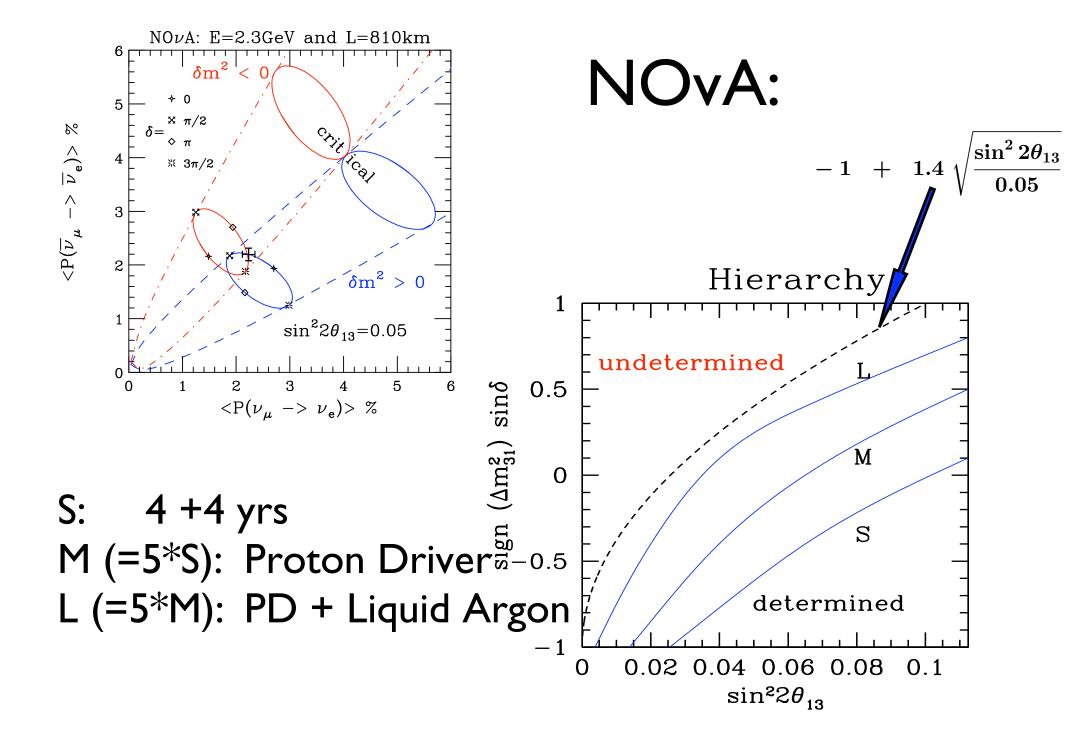
in the overlap region

$$\langle \sin \delta
angle_+ - \langle \sin \delta
angle_- \ = \ 2 \langle heta
angle / heta_{crit} \ pprox \ 1.4 \sqrt{rac{\sin^2 2 heta_{13}}{0.05}}$$

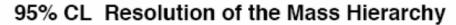
exact along diagonal --- approximately true throughout the overlap region!!!

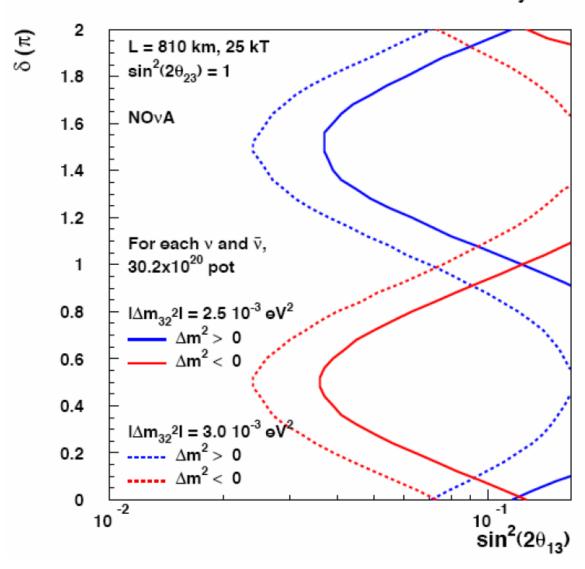
$$heta_{crit} = rac{\pi^2}{8} \; rac{\sin 2 heta_{12}}{\tan heta_{23}} \; rac{\delta m_{21}^2}{\delta m_{31}^2} \left(rac{4\Delta^2/\pi^2}{1-\Delta\cot\Delta}
ight) / (aL) \sim 1/6$$
 i.e. $\sin^2 2 heta_{crit} = 0.10$

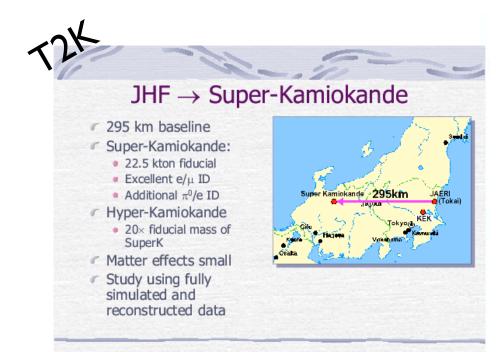
O. Mena + SP hep-ph/0408070

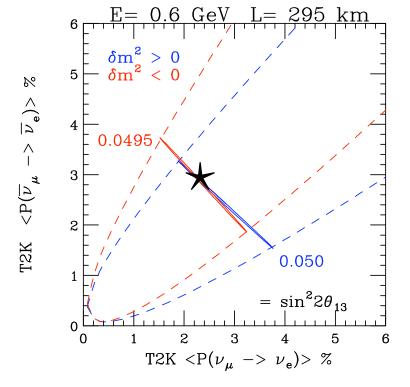


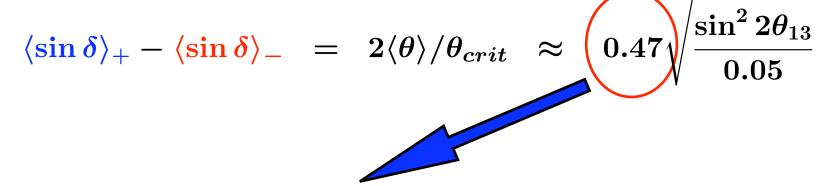
NOvA:



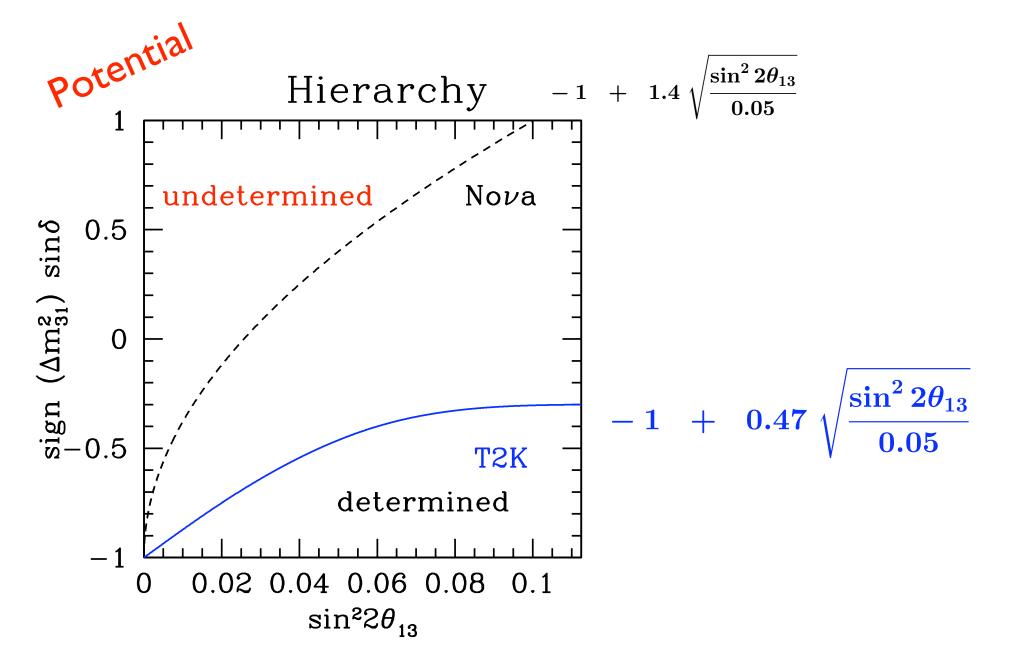




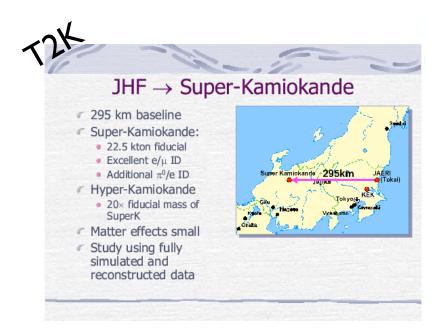


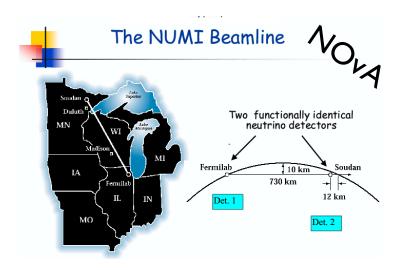


 (ρL) for NOvA three times larger than (ρL) than T2K.



Neutrino v Neutrino Two Expts Different L's and EQUAL <E>/L 's



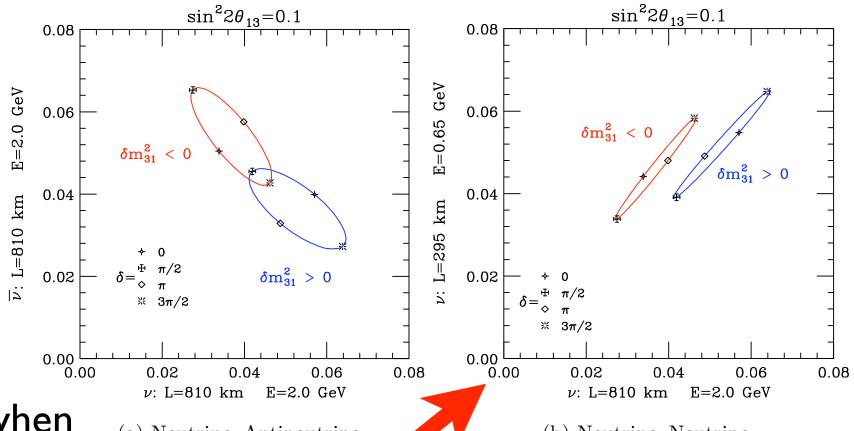


2.5 degrees and I4 kmor2.0 degrees and I2 km

EQUAL <E>/L but NOT

2.5 degrees and I2 km

Bi-Probability nu-antinu v nu-nu



flat when

$$\Delta + \bar{\Delta} = \pi$$

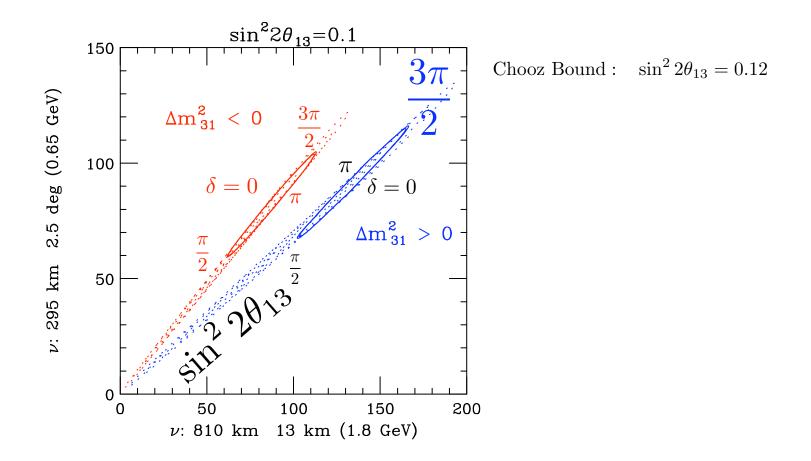
(a) Neutrino–Antineutrino

(b) Neutrino–Neutrino

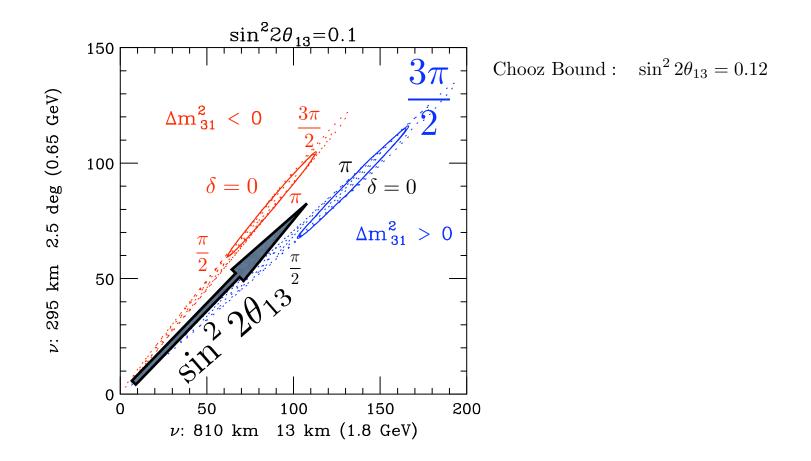
$$P_N(\theta_{13}, \delta) = P_N(\theta_{13}, \delta')$$
 and $P_F(\theta_{13}, \delta) = P_F(\theta_{13}, \delta')$

flat when

non-trivial solutions $\Delta_N = \Delta_F$, same $\langle E \rangle / L$

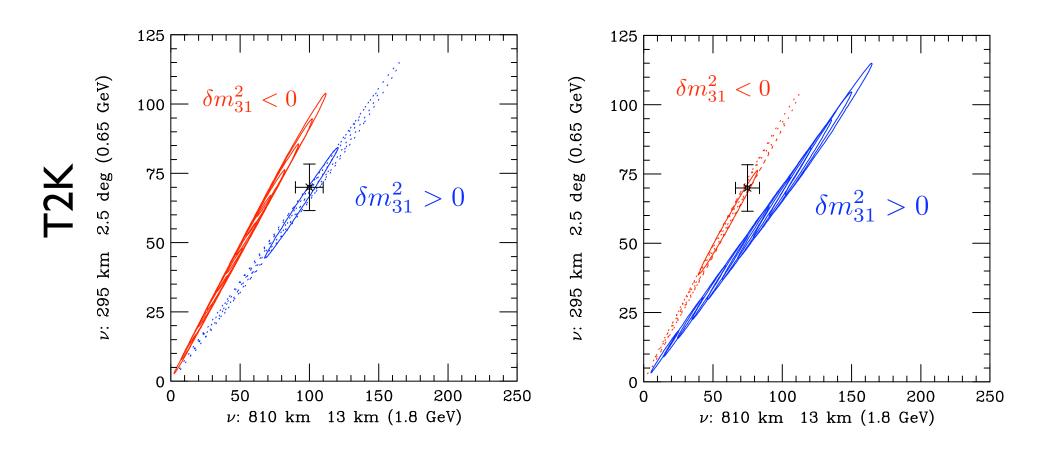


Horiz. separation matter effect for NOvA. Vert. separation matter effect for T2K.



Horiz. separation matter effect for NOvA. Vert. separation matter effect for T2K.

$$\sin^2 2\theta_{13} = 0.07$$



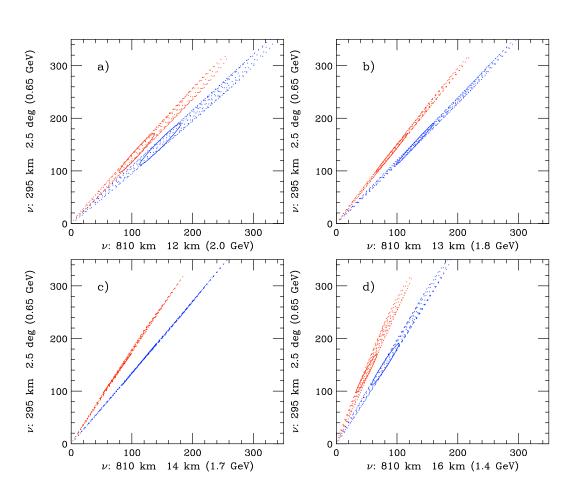
NOvA

NOvA + T2K Neutrinos Only:

Bi-Event Plots:

5 yrs

T2K: 750MW 22.5kton eff=70%



$$\delta m_{31}^2 > 0$$

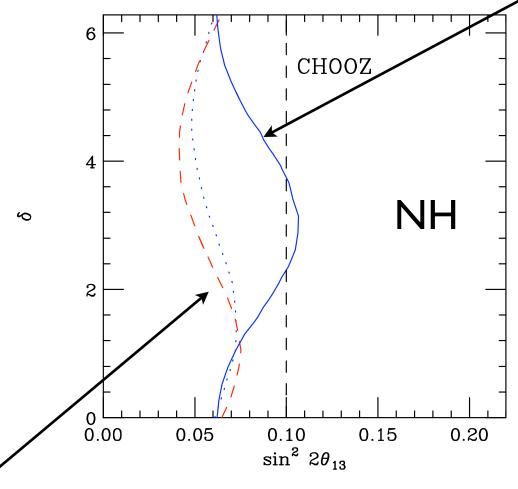
$$\delta m_{31}^2 > 0$$
 $\delta m_{31}^2 < 0$

15% stat. hit per step!

NOvA: 400MW 30kton eff=24%

Hierarchy Determination using 2 dof: 90% CL

2.5 deg & 12 km

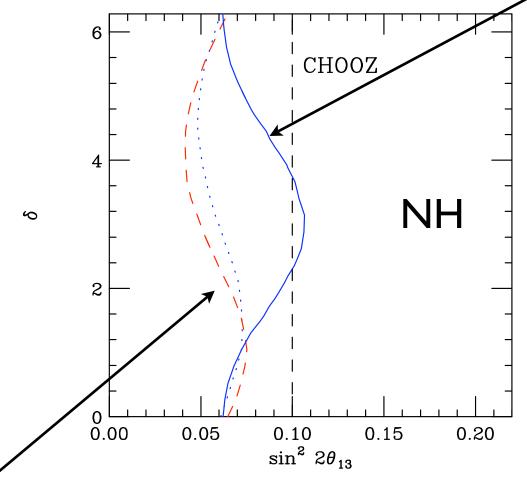


WHY 2 dof????
90% at 2 dof
approx
97% at 1 dof

2.5 deg & I4 km (b) $\Delta m_{31}^2 = +3.0 \times 10^{-3} \text{ eV}^2$ **2.0** deg & I2 km

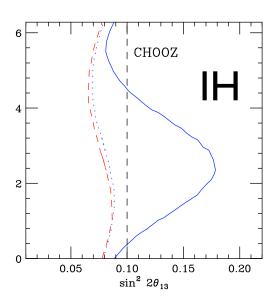
Hierarchy Determination using 2 dof: 90% CL

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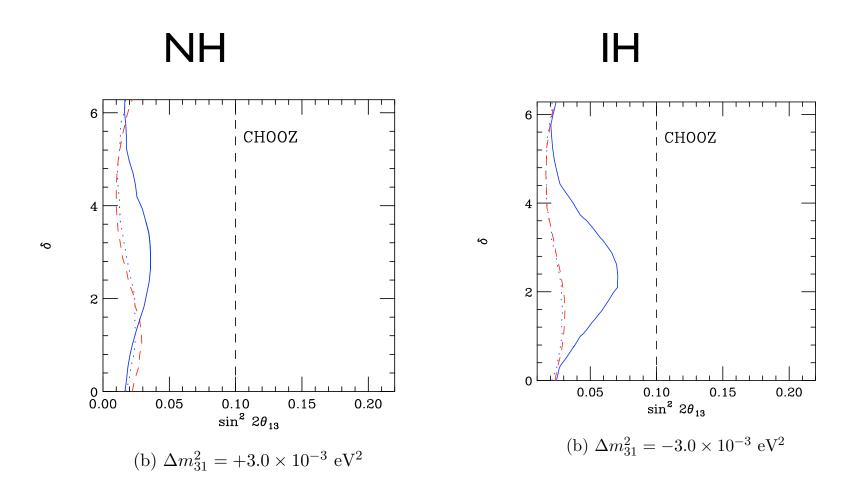
2.5 deg & I4 km (b) $\Delta m_{31}^2 = +3.0 \times 10^{-3} \text{ eV}^2$ **2.0** deg & I2 km

WHY 2 dof????
90% at 2 dof
approx
97% at 1 dof



(b)
$$\Delta m_{31}^2 = -3.0 \times 10^{-3} \text{ eV}^2$$

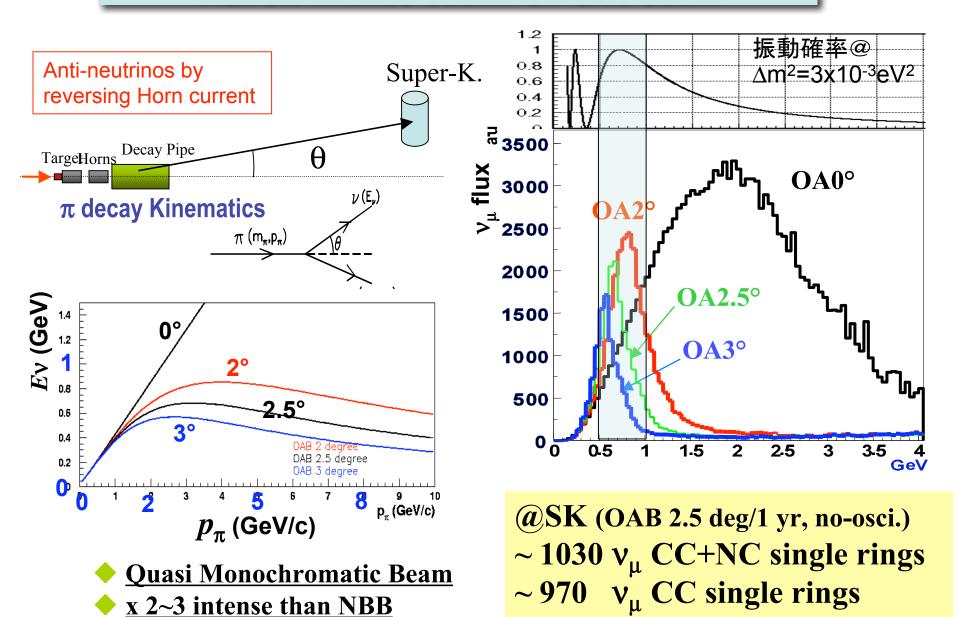
5 times Statistics !!!



systematic errors ???

Aihara for T2K, P5 talk

Narrow intense beam: Off-axis beam



Tuned at oscillation maximum

 $v_e \sim 0.4\%$ at v_μ peak

Suppose T2K ran @ 2.0 deg

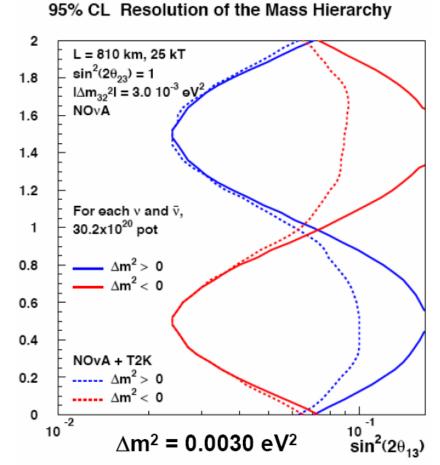
$$\langle E \rangle = 0.75 \text{ MeV}$$

same $\langle E \rangle / L$ as NOvA 12km

- Increased Statistics:
- Improved sensitivity to $\sin^2 \theta_{13}$??? (flux, cross section and backgrounds)
- smaller matter effects:

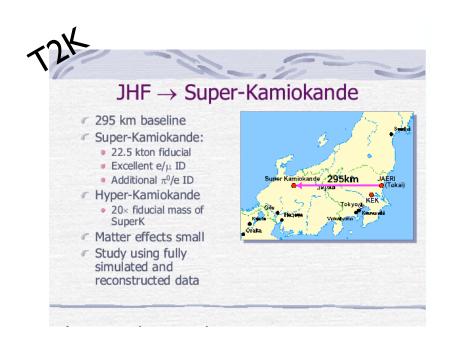
NOvA + T2K:

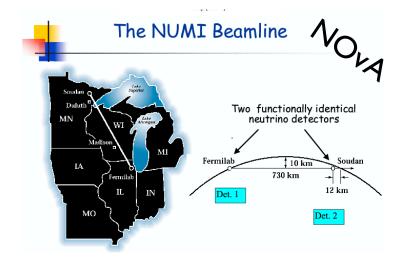
NOvA @ NO-VE 2007

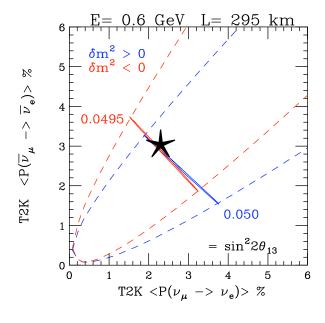


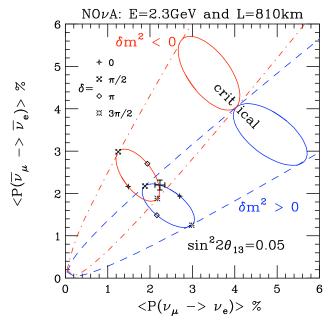
Half Neutrino Running + different <E>/L

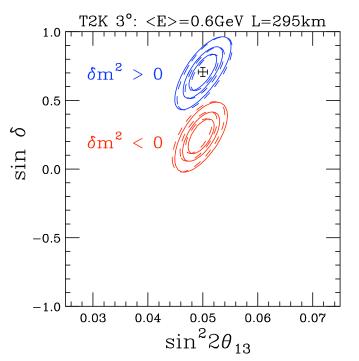
Neutrino v Anti-Neutrino Two Expts. Different L's

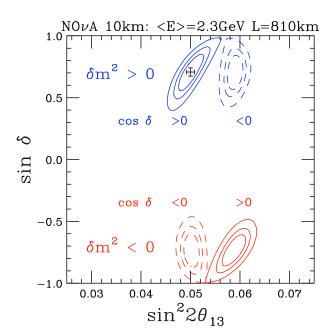












$$egin{array}{lll} \langle \sin \delta
angle_+ & - & \langle \sin \delta
angle_- \ & pprox & 0.47 \sqrt{rac{\sin^2 2 heta_{13}}{0.05}} \end{array}$$

$$\langle \sin \delta
angle_+ - \langle \sin \delta
angle_- \ pprox 1.4 \sqrt{rac{\sin^2 2 heta_{13}}{0.05}}$$

$$|\langle \sin \delta \rangle_{true}^{T2K} - \langle \sin \delta \rangle_{true}^{NO\nu A}| \approx 0$$

$$|\langle \sin \delta \rangle_{fake}^{T2K} - \langle \sin \delta \rangle_{fake}^{NO\nu A}| \approx 0.93 \sqrt{\frac{\sin^2 2\theta_{13}}{0.05}}$$

if the measurement uncertainty on $\sin\delta$

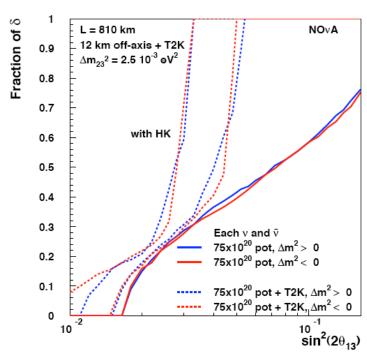
$$\approx \pm 0.2$$

then the two fake solutions are well separated down to

$$\sin^2 2\theta_{13} \approx 0.01$$

Hierarchy is Determined

95% CL



NOvA/PD with T2K Phase 2

Spectrum Measurements:

On Axis

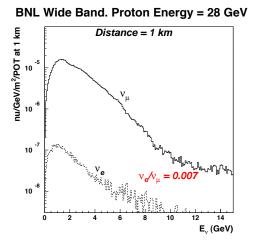
Off Axis - 2nd Peak

On Axis Beams:



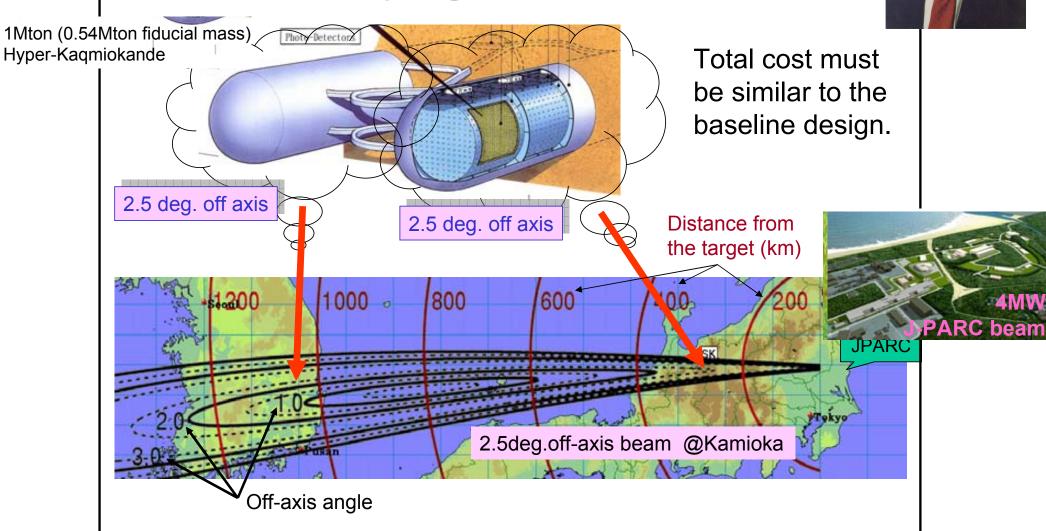
- 28 GeV protons. I MW beam power. Horn focussed
- 500 kT water Cherenkov detector.
- baseline > 2500 km. WIPP, Henderson, Homestake

Brookhaven Proposal



Off Axis:

Some recent progress: detector in Korea



see Kajita talk:

CONCLUSION:

- Determination of Neutrino Hierarchy is Challenging and can effect whether or not the observation of CP violation can be CLAIMED.
- NOvA has a shot, BUT it's VERY δ_{CP} DEPENDENT
- T2K + NOvA can substantially reduce the δ_{CP} dependence using NEUTRINOS ONLY. IF they run at the SAME < E > /L.
 - However, many issues need experimentalist expertise.
 - IMHO we need, join task force between T2K + NOvA before this opportunity is LOST FOREVER !!!